

Starred questions require using computing resources – this isn't a requirement to pass the course, but will help your understanding if you try them.

1. Use the SEMF to predict the *atomic* masses of ^{14}O , and ^{107}Ag , in atomic mass units. Compare these to the real values and estimate the percentage error in each case.

Solution:

^{14}O : predicted: 14.00925965u, real: 14.008596u, error 0.0047%

^{107}Ag : predicted: 106.8974466u, real: 106.905091 u, error 0.0071%

Very accurate!

2. (a) Starting from the SEMF for nuclear masses $M(Z, A)$, rearrange the formula to write it in the form

$$M(Z, A) = \alpha - \beta Z + \gamma Z^2 - \delta/c^2$$

Find expressions for the coefficients α , β and γ .

- (b) Considering the formula as a function of Z for A constant, and neglecting δ , find a formula for $Z(A)$ for the case when the mass is a minimum (ignoring the fact that Z, A are actually integers). Show that for small A this predicts $Z \simeq A/2$. What do you predict for large A ?
- (c) * Using the formula for $Z(A)$ found above make a plot [using mathematica or an equivalent program] of $B(Z, A)/A$ for A up to 240. Comment on any discrepancies between the plot you have done and the plots given in the lecture slides.
- (d) * Make a 3d plot of $B(Z, A)/A$ as a function of N, Z and compare it to the table of nuclides. Discuss your findings.

Solution:

(Note the correction on the δ term from the version you worked from - this didn't affect the question.)

a)

$$\alpha = a_S A^{2/3} + A(a_A - a_V) + m_n, \quad \beta = \frac{a_c}{A^{1/3}} + 4a_A - m_p + m_n, \quad \gamma = \frac{a_c}{A^{1/3}} + 4\frac{a_A}{A}$$

b)

$$Z_{\min} = \frac{4 A^{4/3} a_A + a_c A}{8 a_A A^{1/3} + 2 a_c A}$$

For large A we have $2a_A A^{1/3} / a_C$

Figures shown in the lecture slides of week 3. The main difference is for low A , prediction is a bit high. At high A there are bumps in the real curve not contained in the SEMF.

3. For nuclei with $(N, Z) = (50, 28), (82, 50)$ (i.e., 'doubly magic' nuclei) calculate the binding energies in two ways: a) from the definition of B and by looking up atomic masses for the nuclei, and b) using the SEMF. What do you notice about your answers?

Solution:

$(N, Z) = (50, 28), (82, 50), (126, 82)$ has $B/A = 8.11, 8.275$ MeV. The actual values are: 8.23, 8.35 MeV which are both significantly higher than the SEMF predicts.

4. Use the SEMF to predict the *nuclear* masses of ${}^4\text{He}$, ${}^{14}\text{C}$, ${}^{196}\text{Au}$, ${}^{244}\text{Pu}$ in atomic mass units. Compare these to the real values and estimate the percentage error in each case.

Solution: The (nuclear, atomic, real atomic, % error) predictions are:

${}^4_2\text{He}$ (4.006177186, 4.007274252, 4.00260325413, 0.1166990e-0)

${}^{14}_6\text{C}$ (13.99957648, 14.00286768, 14.003241, 0.266596e-2)

${}^{196}\text{Au}$ (195.9175447, 195.9608788, 195.966569, 0.00290366)

${}^{244}\text{Pu}$ (243.9973552, 244.0489173, 244.064205, 0.626381e-2)

5. In the notes we justified the asymmetry term using evenly spaced energy levels and the Pauli Exclusion Principle. Review this argument, and justify the cumulative effect sequence 1, 2, 5, 8, 13, 18, 25, ... Then show that the energy change must be $\propto (N - Z)^2$. [You can look up on wikipedia a derivation of this based on modelling the nucleus as a Fermi ball of protons and neutrons.]

Solution: This is covered in the notes.

6. Roughly sketch the distribution of mass estimated from the SEMF as a function of Z for $A = 125$ and $A = 128$ around their minima $Z \simeq 52$ and 54 respectively.

Solution: see Krane Fig 3.18

7. Consider isotopes of ${}_{56}\text{Ba}$, with $72 \leq N \leq 88$. Look up data for the neutron separation energy and plot this on a graph as a function of N . Describe the features you see with reference to the SEMF. What feature can't the SEMF explain?

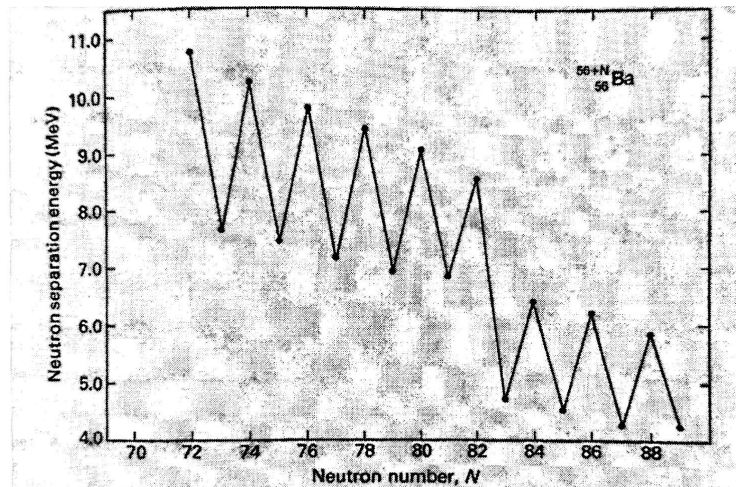


Fig. 4.4 The energy required to remove the last neutron (the neutron separation energy) from the isotopes of barium (${}_{56}\text{Ba}$). This energy is about 2 MeV greater when N is even than when N is odd. This 2 MeV represents the energy required to break the pairing favoured for like nucleons. The irregularity at $N=82$ is one of the effects that we leave for discussion in Chapter 8. (Source of data: Wapstra and Audi 1985.)

Solution:

From Williams' book.

Some (potentially) useful information: The radius of a nuclei may be approximated by $R \approx 1.2A^{1/3}$ fm. The semi-empirical mass formula (SEMF) for the binding energy of a nucleon is

$$B(Z, A) = a_V A - a_S A^{2/3} - a_C \frac{Z(Z-1)}{A^{1/3}} - a_A \frac{(A-2Z)^2}{A} + \delta(Z, A).$$

Constants in the SEMF: $a_V = 15.56, a_S = 17.23, a_C = 0.697, a_A = 23.28, a_p = 12.0$ where each number is in MeV.

Nuclear Shells: Protons

$$1s_{\frac{1}{2}} \downarrow_2 \quad 1p_{\frac{3}{2}} \downarrow_4 \quad 1p_{\frac{1}{2}} \downarrow_2 \quad 1d_{\frac{5}{2}} \downarrow_6 \quad 2s_{\frac{1}{2}} \downarrow_2 \quad 1d_{\frac{3}{2}} \downarrow_4 \quad 1f_{\frac{7}{2}} \downarrow_8 \quad 2p_{\frac{3}{2}} \downarrow_4 \quad 1f_{\frac{5}{2}} \downarrow_6 \quad 2p_{\frac{1}{2}} \downarrow_2 \quad 1g_{\frac{9}{2}} \downarrow_{10} \quad 1g_{\frac{7}{2}} \downarrow_8 \quad 2d_{\frac{5}{2}} \downarrow_6 \quad 1h_{\frac{11}{2}} \downarrow_{10} \quad 2d_{\frac{3}{2}} \downarrow_4 \quad 3s_{\frac{1}{2}} \downarrow_2 \quad 1h_{\frac{9}{2}} \downarrow_8 \quad 2f_{\frac{7}{2}} \downarrow_6 \quad \dots$$

Shells: Neutrons

$$1s_{\frac{1}{2}} \downarrow_2 \quad 1p_{\frac{3}{2}} \downarrow_4 \quad 1p_{\frac{1}{2}} \downarrow_2 \quad 1d_{\frac{5}{2}} \downarrow_6 \quad 2s_{\frac{1}{2}} \downarrow_2 \quad 1d_{\frac{3}{2}} \downarrow_4 \quad 1f_{\frac{7}{2}} \downarrow_8 \quad 2p_{\frac{3}{2}} \downarrow_4 \quad 1f_{\frac{5}{2}} \downarrow_6 \quad 2p_{\frac{1}{2}} \downarrow_2 \quad 1g_{\frac{9}{2}} \downarrow_{10} \quad 2d_{\frac{5}{2}} \downarrow_6 \quad 1g_{\frac{7}{2}} \downarrow_8 \quad 1h_{\frac{11}{2}} \downarrow_{10} \quad 2d_{\frac{3}{2}} \downarrow_4 \quad 3s_{\frac{1}{2}} \downarrow_2 \quad 2f_{\frac{7}{2}} \downarrow_6 \quad 1h_{\frac{9}{2}} \downarrow_8 \quad \dots$$

$\frac{e^2}{4\pi\epsilon_0}$	$= 1.439965 \text{ MeV fm}$
Boltzmann's constant	$k_B = 8.6173303 \times 10^{-5} \text{ eV/K}$
Planck's constant	$h = 4.135668 \times 10^{-15} \text{ eV s}$
Speed of light	$c = 2.99792 \times 10^8 \text{ m/s}$
Neutrino mean lifetime	881 s
Atomic mass unit	$1 u = 931.4940954 \text{ MeV}/c^2 = 1.66054 \times 10^{-27} \text{ kg}$
Mass of electron	$m_e = 5.4858 \times 10^{-4} u = 0.51099895 \text{ MeV}/c^2$
Mass of proton	$m_p = 1.00727646688 u = 938.27208 \text{ MeV}/c^2$
Mass of neutron	$m_n = 1.00866491578 u = 939.56541 \text{ MeV}/c^2$
Mass of ^1_1H	$= 1.00782503 u$
Mass of ^2_1H	$= 2.01410178 u$
Mass of ^3_1H	$= 3.01604927 u$
Mass of ^3_2He	$= 3.01602932 u$
Mass of ^4_2He	$= 4.00260325 u$
Mass of $^{232}_{90}\text{Th}$	$= 232.038055 u$
Mass of $^{234}_{90}\text{Th}$	$= 234.043601 u$
Mass of $^{235}_{92}\text{U}$	$= 235.043930 u$
Mass of $^{236}_{92}\text{U}$	$= 236.045568 u$
Mass of $^{238}_{92}\text{U}$	$= 238.050788 u$
Mass of $^{239}_{92}\text{U}$	$= 239.054293 u$
Mass of $^{240}_{94}\text{Pu}$	$= 240.053811 u$
Mass of $^{241}_{94}\text{Pu}$	$= 241.056849 u$
Mass of $^{242}_{94}\text{Pu}$	$= 242.058741 u$
Mass of the Sun	$M_\odot = 1.988 \times 10^{30} \text{ kg}$
Gravitational constant	$G = 6.67408 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$

Nuclei masses given are atomic masses.

You can look up other nuclear data from websites

<https://www-nds.iaea.org/relnsd/vcharthtml/VChartHTML.html>

<http://www.nndc.bnl.gov/nudat2/>

<http://atom.kaeri.re.kr/nuchart/>

<http://people.physics.anu.edu.au/~ecs103/chart/>